

HIP-1999-45/TH
KANAZAWA-99-16
July, 1999

***F*-term Inflation in M-theory with Five-branes**

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Abstract

We study *F*-term inflation in M-theory with and without five-brane moduli fields. We show the slow rolling condition is not satisfied in M-theory without five-brane moduli fields, but it can be satisfied in the case with non-vanishing *F*-terms of five-brane moduli fields.

1 Introduction

Cosmological inflation can solve the flatness and horizon problems of the universe [1]. During inflation, the vacuum energy $V_0 \equiv \langle V \rangle$ takes a value $V_0 = 3H^2 M^2$, where M is the reduced Planck scale and H is expected to be of $O(10^{13} \sim 10^{14})$ GeV. The successful inflation scenario also requires the so-called slow-rolling condition for the inflaton field Φ , that is, the parameters $\varepsilon_{in} = \frac{1}{2}M^2(V'/V)^2$ and $\eta = M^2 V''/V$ should be suppressed,

$$\varepsilon_{in} \ll 1, \quad \eta \ll 1. \quad (1)$$

Within the framework of supergravity, the scalar potential V includes the F -term and D -term contributions,

$$V = F^I \bar{F}^{\bar{J}} \partial_I \partial_{\bar{J}} K - 3e^G M^4 + (D\text{-term}). \quad (2)$$

Here K denotes the Kähler potential and G is obtained as $G = KM^{-2} + \ln(|W|^2 M^{-6})$, where W is the superpotential. Thus, the nonvanishing vacuum energy V_0 can be induced by nonvanishing F -terms and/or D -terms. That implies supersymmetry breaking.

Here we consider the case that F -terms contributions are dominant within the framework of string-inspired supergravity [2, 3, 4], although the D -term inflation [5] induced by the anomalous $U(1)$ symmetry [6] is also another interesting possibility. Within the framework of superstring theory, the dilaton and moduli fields can be candidates for the inflation field [2]. However, here we consider other matter fields as candidates for the inflaton in hybrid inflation, which is driven by the vacuum energy due to non-vanishing F -terms of the dilaton and moduli fields.

In general, non-vanishing F -terms, however, induce a seizable soft scalar mass m_Φ for the inflaton field Φ , and that spoils the flatness condition (1) of V for Φ . This problem has been discussed within the framework of weakly coupled superstring theory. In Ref.[7] it has been shown that we have $m_\Phi^2 = 0$ for the field with the modular weight $n = -3$ and the vacuum energy V_0 driven by the F -term of the moduli field T in weakly coupled heterotic string theory without nonperturbative Kähler potential.

In this paper we consider the condition $m_\Phi^2 = 0$ for $V_0 \neq 0$ within the framework of strongly coupled heterotic string theory, M-theory [8]. We take into account effects due to five-brane moduli fields.

2 M-theory without five-brane

First we consider the case without five-brane. The Kähler potential K is obtained [9],

$$K = -\log(S + \bar{S}) - 3\log(T + \bar{T}) + \left(\frac{3}{T + \bar{T}} + \frac{\alpha}{S + \bar{S}} \right) |\Phi|^2, \quad (3)$$

where $\alpha = 1/(8\pi^2) \int \omega \wedge [tr(F \wedge F) - \frac{1}{2}tr(R \wedge R)]$. The fields S and T denote the dilation field and moduli field. We assume that the F -terms of S and T contribute to V_0 . Then we parameterize the F -terms,

$$F^S = \sqrt{3}m_{3/2}C(S + \bar{S})\sin\theta, \quad (4)$$

$$F^T = m_{3/2}C(T + \bar{T})\cos\theta, \quad (5)$$

where $C^2 = 1 + V_0/(3m_{3/2}^2M^2)$ and $m_{3/2} = e^G M^2$. We have neglected the CP phases of F -terms. Using this parametrization, the soft scalar mass m_Φ^2 is obtained [10],

$$\begin{aligned} m_\Phi^2 = & V_0 M^{-2} + m_{3/2}^2 - \frac{3C^2 m_{3/2}^2}{3(S + \bar{S}) + \alpha(T + \bar{T})} \\ & \times \left\{ \alpha(T + \bar{T}) \left(2 - \frac{\alpha(T + \bar{T})}{3(S + \bar{S}) + \alpha(T + \bar{T})} \right) \sin^2\theta \right. \\ & + (S + \bar{S}) \left(2 - \frac{3(S + \bar{S})}{3(S + \bar{S}) + \alpha(T + \bar{T})} \right) \cos^2\theta \\ & \left. - \frac{2\sqrt{3}\alpha(T + \bar{T})(S + \bar{S})}{3(S + \bar{S}) + \alpha(T + \bar{T})} \sin\theta \cos\theta \right\}. \end{aligned} \quad (6)$$

We take the limit $C^2 \gg 1$, i.e. $V_0 \gg 3m_{3/2}^2 M^2$, and investigate the condition for $m_\Phi^2 \ll O(H^2 = V_0 M^{-2})$. It is obvious that we have $m_\Phi^2 = O(H^2)$ in most of the parameter space. However, the condition $m_\Phi^2 = 0$ is realized for the following case:

$$\cos\theta = 0, \quad \frac{(S + \bar{S})}{\alpha(T + \bar{T})} = 0. \quad (7)$$

Unfortunately, the solution (7) is not a realistic solution from the viewpoint of M-theory. In M-theory, we have two sectors, which are usually called

the observable sector and the hidden sector. The sector including the scalar field Φ has the gauge kinetic function f [11],

$$f = S + \alpha T, \quad (8)$$

and the other sector has the gauge kinetic function f' ,

$$f' = S + \alpha' T. \quad (9)$$

The gauge couplings of these sectors g and g' are obtained as $g^{-2} = \text{Re}(f)$ and $g'^{-2} = \text{Re}(f')$. The coefficients α and α' should satisfy

$$\alpha + \alpha' = 0. \quad (10)$$

It is impossible that the both of g^{-2} and g'^{-2} take positive values for the condition (7), i.e. $(S + \bar{S}) \ll \alpha(T + \bar{T})$. Thus, the condition $m_\Phi^2 \ll O(H^2)$ can not be realized in the F -term inflation of M-theory without five-brane.

3 M-theory with five-branes

Next we consider M-theory with five-brane moduli fields Z^n [12], whose vacuum expectation values provide with positions z^n of the corresponding five-branes along the orbifold S^1/Z_2 . The moduli Kähler potential K_{mod} and the Kähler potential of the scalar field Φ K_Φ are obtained

$$K_{mod} = -\ln(S + \bar{S}) - 3\ln(T + \bar{T}) + K_5, \quad (11)$$

$$K_\Phi = \left(\frac{3}{(T + \bar{T})} + \frac{\epsilon \zeta}{(S + \bar{S})} \right) |\Phi|^2, \quad (12)$$

where

$$\zeta = \beta + \sum_{n=1}^N (1 - z^n)^2 \beta^{(n)}. \quad (13)$$

Here K_5 denotes the Kähler potential for Z^n , ϵ is the expansion parameter $\epsilon = (\kappa/4\pi)^{2/3} 2\pi^2 \rho/V^{2/3}$, $\beta^{(n)}$ is a magnetic charge on the each 5-brane and β is the instanton number on the sector boundary including the inflaton Φ . The gauge kinetic functions of the sector including the inflaton Φ and the

other sector, f and f' , are obtained in the same forms as eqs.(8) and (9) except changing α and α' ,

$$\begin{aligned}\alpha &= \epsilon \left(\beta + \sum_{n=1}^N (1 - z^n)^2 \beta^{(n)} \right), \\ \alpha' &= \epsilon \left(\beta' + \sum_{n=1}^N (z^n)^2 \beta^{(n)} \right).\end{aligned}\tag{14}$$

We have the constraint

$$\beta + \beta' + \sum_{n=1}^N \beta^{(n)} = 0.\tag{15}$$

The condition (10) is relaxed by $\beta^{(n)}$. Thus the solution (7) could be realistic and both g^{-2} and g'^{-2} could be positive if both α and α' are positive, or if one of them is positive and the other is suppressed enough, e.g. $\alpha > 0$ and $\alpha' \approx 0$.

Furthermore, the F -term of the five-brane moduli Z^n can also contribute to V_0 and m_Φ^2 . That changes the situation and could give a realistic solution even for $\alpha\alpha' < 0$. In this section, we consider effects of the F -term of Z^n .

For simplicity, we assume that there is only one relevant 5-brane moduli Z and its Kähler potential is a function of only $(Z + \bar{Z})$. Now we can parameterize each F -term as follows

$$\begin{aligned}F^S &= \sqrt{3}m_{3/2}C(S + \bar{S})\sin\theta\sin\phi, \\ F^T &= m_{3/2}C(T + \bar{T})\cos\theta\sin\phi, \\ F^Z &= \sqrt{\frac{3}{K_{5,Z\bar{Z}}}}m_{3/2}C\cos\phi.\end{aligned}\tag{16}$$

In addition, we fix z and $\beta^{(1)}$, e.g.

$$z = \frac{1}{2}, \quad \beta^{(1)} = \frac{4}{3}\beta.\tag{17}$$

Note that in this case we have

$$2\text{Re}(f) = (S + \bar{S}) + \epsilon\zeta(T + \bar{T}),\tag{18}$$

$$2\text{Re}(f') = (S + \bar{S}) - \frac{3}{2}\epsilon\zeta(T + \bar{T}),\tag{19}$$

i.e. $\alpha\alpha' < 0$, where $\alpha = \epsilon\zeta$ and $\alpha' = -3\epsilon\zeta/2$.

In the limit $C \gg 1$, we obtain m_Φ^2 [13] *,

$$\begin{aligned}
m_\Phi^2 = & 3m_{3/2}^2 C^2 \left\{ 1 - \left[1 - \frac{9(S + \bar{S})^2}{(3(S + \bar{S}) + \epsilon\zeta(T + \bar{T}))^2} \right] \sin^2 \theta \sin^2 \phi \right. \\
& - \frac{1}{3} \left[1 - \left(\frac{\epsilon\zeta(T + \bar{T})}{3(S + \bar{S}) + \epsilon\zeta(T + \bar{T})} \right)^2 \right] \cos^2 \theta \sin^2 \phi \\
& - \frac{\epsilon\zeta(T + \bar{T})}{K_{5,Z\bar{Z}}(3(S + \bar{S}) + \epsilon\zeta(T + \bar{T}))} \left[2 - \frac{\epsilon\zeta(T + \bar{T})}{(3(S + \bar{S}) + \epsilon\zeta(T + \bar{T}))} \right] \cos^2 \phi \\
& + \frac{2\sqrt{3}(S + \bar{S})\epsilon\zeta(T + \bar{T})}{(3(S + \bar{S}) + \epsilon\zeta(T + \bar{T}))^2} \sin \theta \cos \theta \sin^2 \phi \\
& - \frac{2\sqrt{3}(S + \bar{S})\epsilon\zeta(T + \bar{T})}{(3(S + \bar{S}) + \epsilon\zeta(T + \bar{T}))^2} \sqrt{\frac{3}{K_{5,Z\bar{Z}}}} \sin \theta \sin \phi \cos \phi \\
& + \frac{2\epsilon\zeta(T + \bar{T})}{3(S + \bar{S}) + \epsilon\zeta(T + \bar{T})} \left[1 - \frac{\epsilon\zeta(T + \bar{T})}{(3(S + \bar{S}) + \epsilon\zeta(T + \bar{T}))} \right] \\
& \times \left. \sqrt{\frac{1}{3K_{5,Z\bar{Z}}}} \cos \theta \sin \phi \cos \phi \right\}. \tag{20}
\end{aligned}$$

The region with $\cos \theta = 0$ in eq.(6) gives a solution for $m_\Phi^2 = 0$ without five-brane, although it is not realistic because of negative gauge couplings squared. Hence, let us consider the condition $m_\Phi^2 = 0$ for $\cos \theta = 0$ in eq.(20) at first. For $K_{5,Z\bar{Z}} = 1$, the equation becomes simple and the two solutions for $m_\Phi^2 = 0$ are obtained

$$\frac{(S + \bar{S})}{\alpha(T + \bar{T})} = 0, \quad \frac{(S + \bar{S})}{\alpha(T + \bar{T})} = \frac{2}{3} \sin \phi \cos \phi. \tag{21}$$

While the former corresponds to the solution (7) in the case without five-brane, the latter is a new solution. However, the latter is not realistic, either, because the value $\alpha(T + \bar{T})$ is still large compared with $(S + \bar{S})$ and it can not lead to positive g^{-2} and g'^{-2} at the same time.

*See also Ref.[14].

Next, let us consider the $K_{5,Z\bar{Z}}$ dependence varying it for $\cos\theta = 0$. We fix ϕ , e.g. $\tan\phi = 1$. On top of that, the gauge coupling constant of the observable sector g_{GUT} must satisfy $2g_{GUT}^{-2} \simeq 4$. Thus, if the field Φ belongs to the observable sector, we take

$$(S + \bar{S}) + \alpha(T + \bar{T}) = 4. \quad (22)$$

On the other hand, if the field Φ belongs to the hidden sector, we take

$$(S + \bar{S}) - \frac{3}{2}\alpha(T + \bar{T}) = 4. \quad (23)$$

In the former case (22) the positivity of g^{-2} and g'^{-2} requires $\alpha(T + \bar{T}) < 8/5$, while in the latter case (23) $\alpha(T + \bar{T}) > -8/5$ is required. Fig.1 shows the solution of $m_\Phi^2 = 0$ leading to such values of $\tau = \alpha(T + \bar{T})$. The lower and upper lines correspond to the cases (22) and (23), respectively. In the case that Φ belongs to the observable sector, $1/K_{5,Z\bar{Z}} > 3.4$ is required for $\alpha(T + \bar{T}) < 8/5$. In the case that Φ belongs to the hidden sector, the same region $1/K_{5,Z\bar{Z}} > 3.4$ leads to positive values of $\alpha(T + \bar{T})$, that is, $\alpha(T + \bar{T}) > -8/5$ is satisfied. In this region, we have the realistic solutions of $m_\Phi^2 = 0$ for $\cos\theta = 0$ and $\tan\phi = 1$ when the field Φ belongs to the observable or hidden sector.

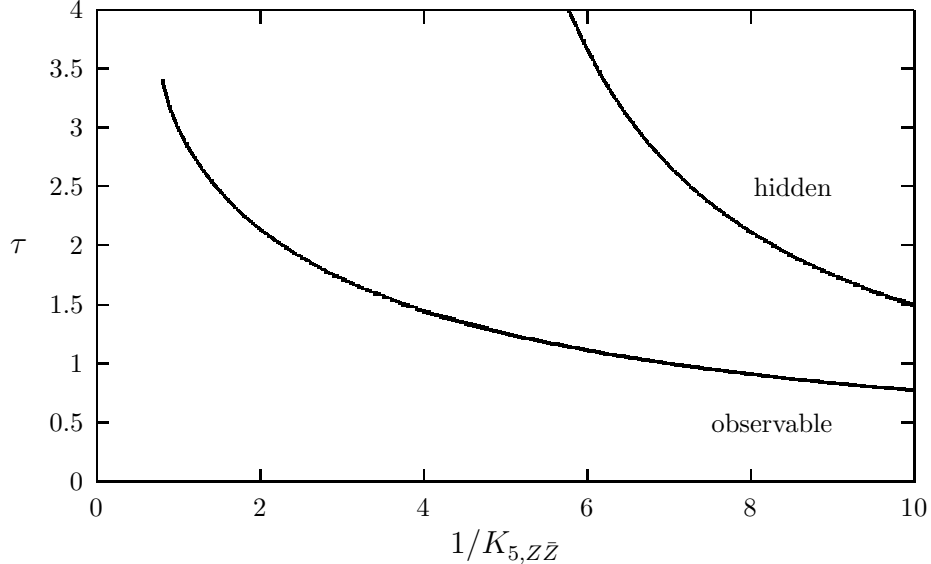


Fig.1: Solutions of $m_\Phi^2 = 0$ for $\cos\theta = 0$.

Similarly, we can investigate the condition $m_\Phi^2 = 0$ for other values of θ . Let us discuss the case with $\tan \theta = 1/\sqrt{3}$ as another example. We concentrate to the case that Φ belongs to the observable sector. When we fix $K_{5,Z\bar{Z}} = 1$, we have only the non-realistic solution, $(S + \bar{S})/(\alpha(T + \bar{T})) = 0$, i.e. $\alpha(T + \bar{T}) = 4$. Now let us vary $K_{5,Z\bar{Z}}$ fixing ϕ , e.g. $\tan \phi = 1$. Fig. 2 show the realistic solution against $K_{5,Z\bar{Z}}$ for $\tan \phi = 1$. In this case, the region $1/K_{5,Z\bar{Z}} > 5$ is favorable.

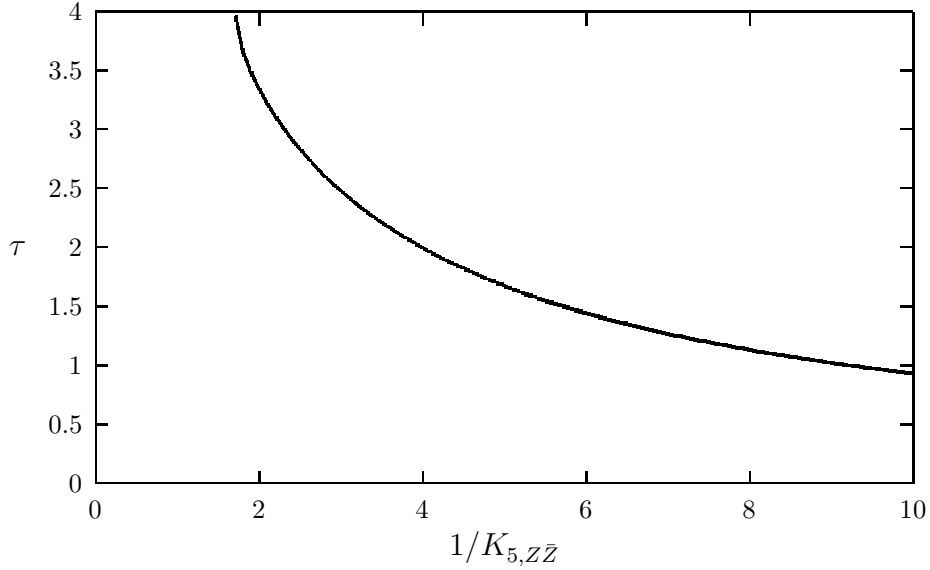


Fig.2: The solution $m_\Phi^2 = 0$ for $\tan \theta = 1$.

4 Conclusion

We have studied on the flatness condition within the framework of M-theory with and without five-brane. In the case without five-brane, we can not obtain a realistic solution, because the condition (10) constrains severely. In the case with five-brane, the F -terms of five-brane moduli fields can also contribute the supersymmetry breaking and the vacuum energy. We have shown such effects are important to obtain the realistic solutions for $m_\Phi^2 = 0$. In particular, a large value of $1/K_{5,Z\bar{Z}}$ is favorable. In addition, the condition (10) is relaxed because of five-brane effects.

Acknowledgement

The authors would like to thank J. Kubo and M. Roos for their encouragements and useful discussions. The research of T.K. was partially supported by the Academy of Finland (no. 44129). S.M.M. is indebted to the Magnus Ehrnrooths Foundation for its support.

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